

Long-term change in the cyclotron line energy in Hercules X-1

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R. Staubert HEAD Fifteenth Div. Meeting, Naples, Florida, 05 April 2016

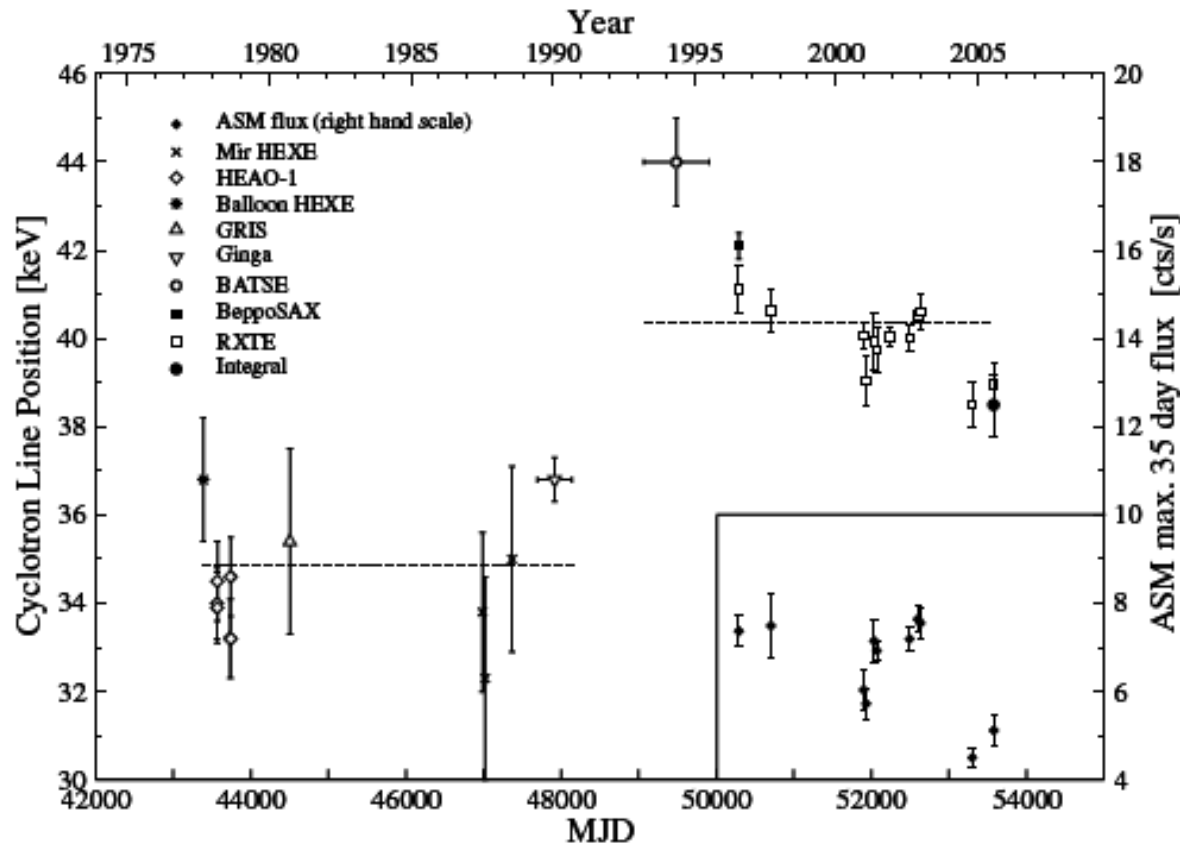
Variability in the Cyclotron Line Energy E_{cyc} in Hercules X-1

E_{cyc} varies with

- **pulse phase** (not discussed here)
- **35d phase** (not discussed here)
- **luminosity**
- **time** (long-term, 10s of years)

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Long-term evolution of E_{cyc}



Her X-1 2007

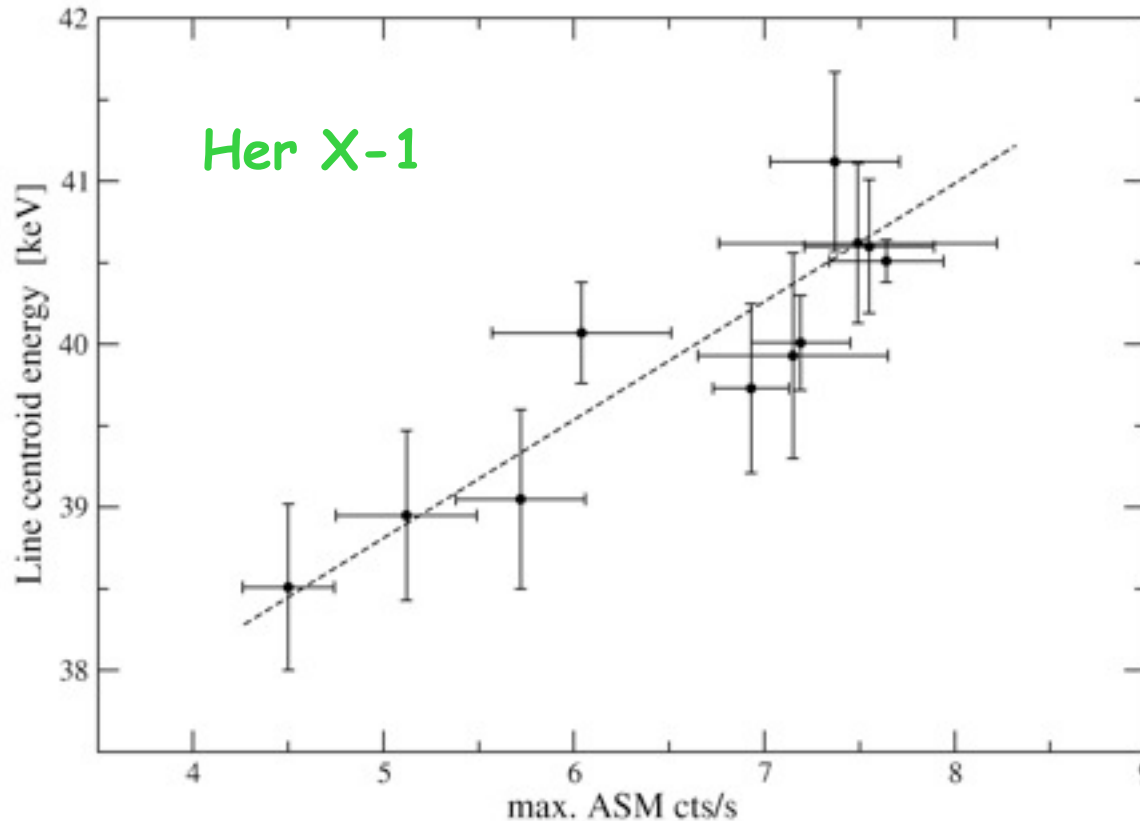
Staubert et al. 2007,
A&A 465, L25

- upward jump in E_{cyc} btw. 1991 and 1993
- apparent decay after 1993: is it real?
- does E_{cy} correlate with flux?

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Positive correlation of E_{cyc} with L_x

- first detection in a persistent source at moderate luminosity



Staubert et al. 2007,
A&A 465, L25

E_{cyc} increases
by ~6% for a factor
of 2 increase in L_x

in V 0332+63
(at high luminosity)
 E_{cyc} is known to
decrease with L_x

observations 1996 / 2006: E_{cyc} [keV] = 40 + 0.66 (max. ASM cts/s - 6.8)

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Objects with E_{cyc} increasing with increasing L_x

- Her X-1** Staubert et al. 2007, A&A 465, L25
- GX 304-1** Yamamoto et al. 2011, PASJ 63, 751
Klochkov et al. 2012, A&A 542, L28
Rothschild et al. 2016 (see poster)
- A 0535+26** Klochkov et al. 2011, A&A 532, A126
- Vela X-1** Fürst et al. 2014, ApJ 780, 133
- Cep X-4** Fürst et al. 2014, ApJ 806, L24

to be confirmed:

- GX 301-2** La Barbera et al. 2005, A&A 438, 617
- 4U 1538-522** Hemphill et al. 2016, MNRAS accepted
(arXiv: 1603.02198)

Only one object with E_{cyc} decreasing with increasing L_x :

V 0332+53

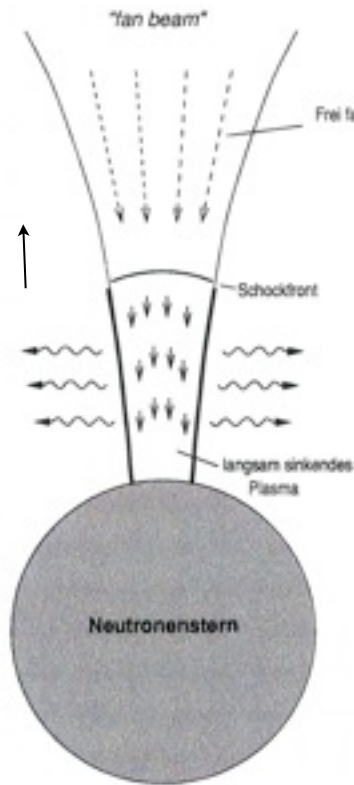
Tsygangov et al. 2006,
MNRAS 371, 19

Klochkov et al. 2011,
A&A 532, A126
(pulse-to-pulse)

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Two accretion regimes

height increases
for increasing L_x



$$L_x > L_{\text{crit}}$$

observations:

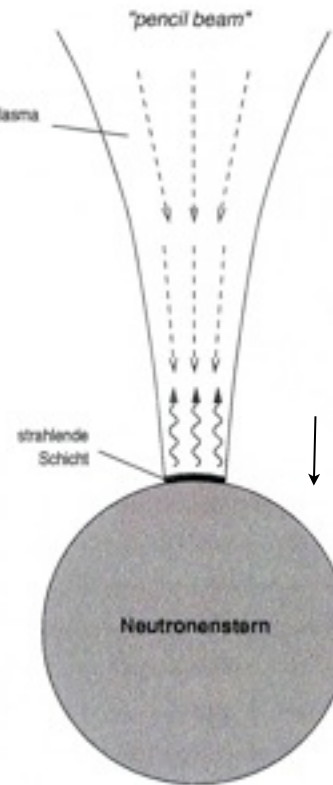
Staubert et al. 2007, *A&A* 465, L25

Klochkov et al. 2011, *A&A* 532, 126

theory:

Becker et al. 2012, *A&A* 544, 123

height decreases
for increasing L_x

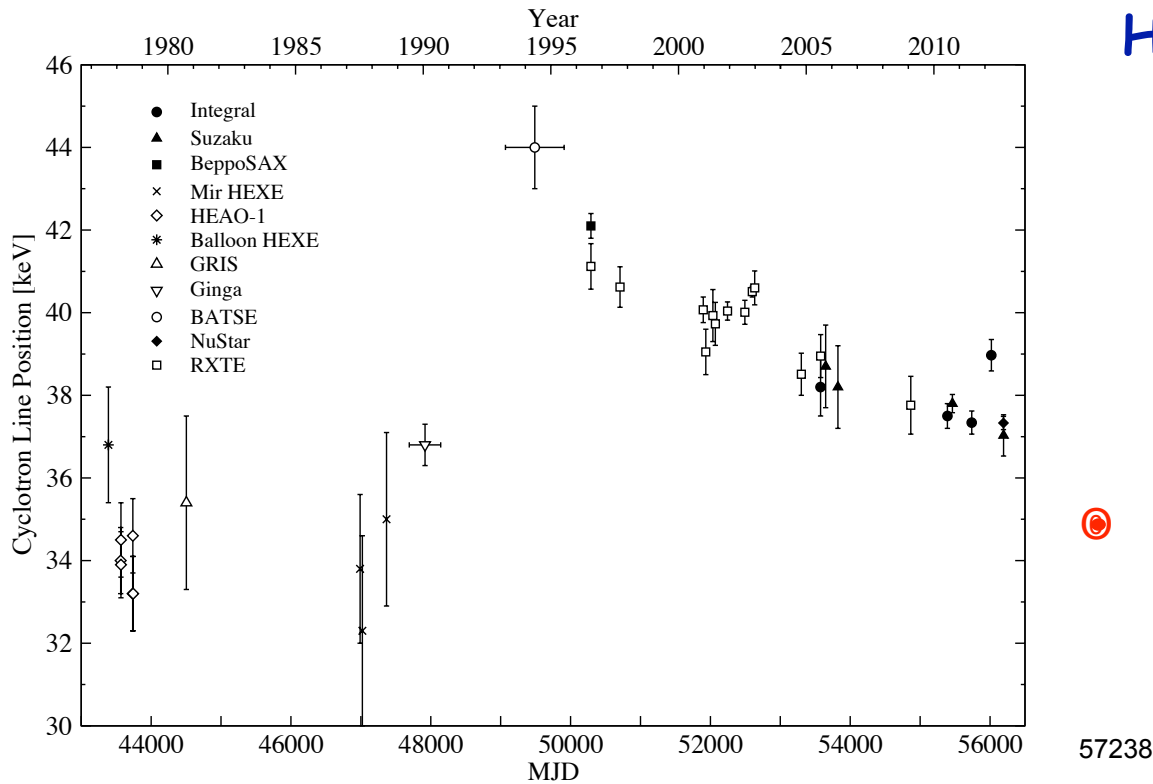


$$L_x < L_{\text{crit}}$$

$$L_{\text{crit}} \sim 10^{37} \text{ erg/s}$$

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Long-term evolution of E_{cyc}



Her X-1 2014

Staubert et al. 2014
A&A 572, A119

Staubert et al. 2016
A&A in print
arXiv: 1603.07090

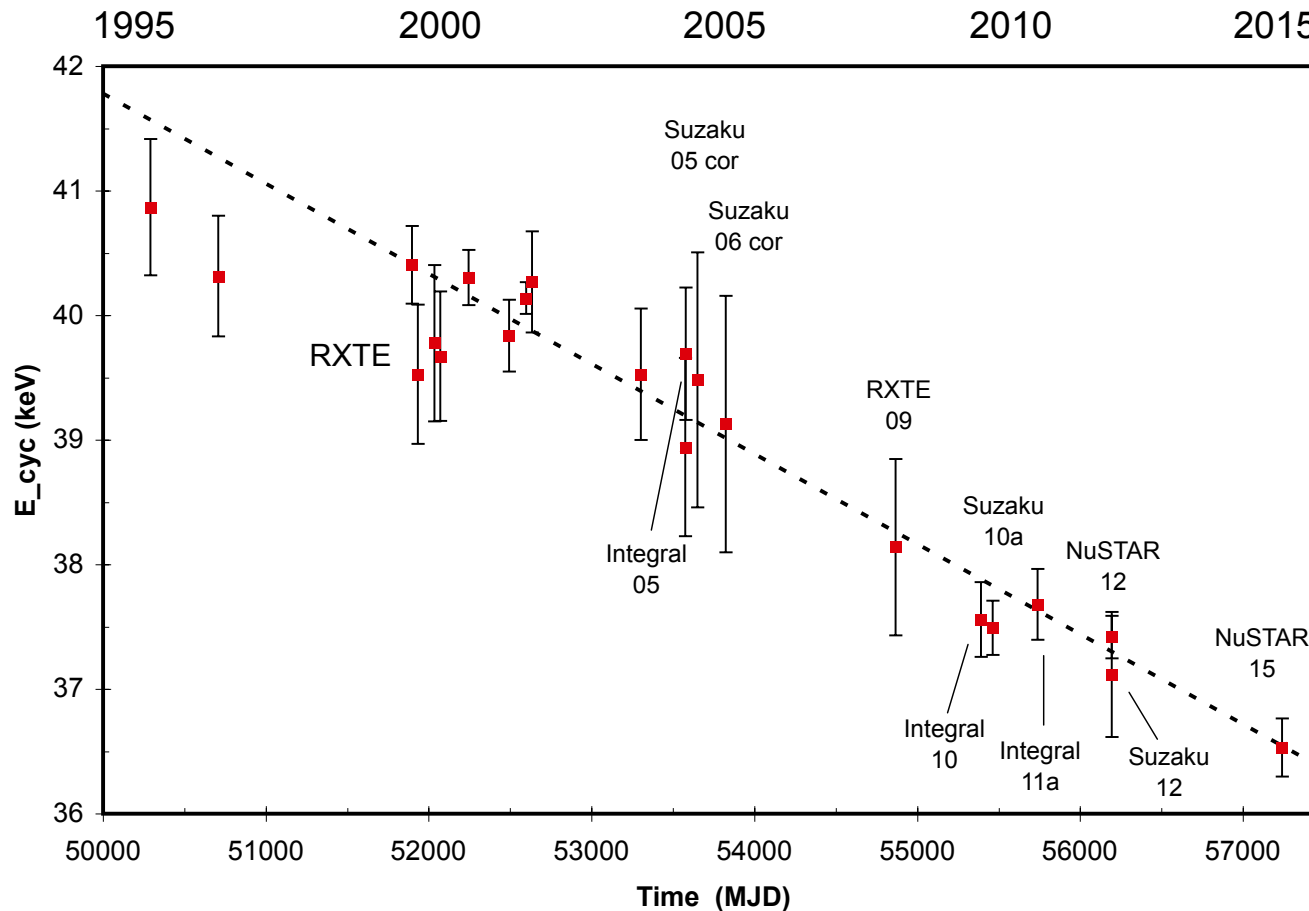


NuSTAR 2015:
34.8 +/- 0.2 keV
at MJD 57238

the decay of E_{cyc} continued !!

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Her X-1: continued decay of E_{cyc} - new NuSTAR point



Her X-1 2016

Staubert et al. 2016

A&A in print

arXiv: 1603.07090

$$E_0 = (39.25 \pm 0.07) \text{ keV}$$

$$a = (0.44 \pm 0.09) \text{ keV}/(\text{ASM-cts/s})$$

$$b = (-7.22 \pm 0.39) 10^{-4} \text{ keV/d}$$

**$b = -5 \text{ keV}$
in 20 yrs**

$F_0 = 6.8 \text{ ASM-cts/s}$

$T_0 = \text{MJD } 53500$

best fit with two variables (Flux and Time): $E_{cyc} = E_0 + a * (F - F_0) + b * (T - T_0)$

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Physics of the E_{cyc} change with time ??

Ideas (comments /questions):

- Change of the global dipole field of the NS?

unlikely

more likely a local phenomenon:

- Accretion mound structure changes due to continuous accretion?

- increase of the height of the accretion mound? dipole: -5 keV means +400 m !?

- hot spot area is increased, B-field is diluted?

- B-field is "screened" or "buried"?

total mass needed? time scale?

- "Ohmic dissipation" of B-field? on "diffusion" time scale:

$t = 4 \pi R^2 \sigma / c^2$, needs small R and small σ !

- Thermo-magnetic effects or Hall effect?

heating due to accretion?

- ???????

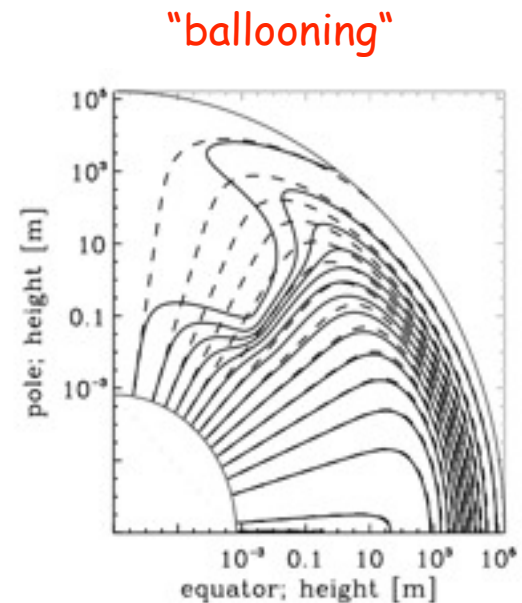
Basic questions about accretion onto NSs

Assume continuous accretion (e.g. 10^{17} g/sec):

- what happens to the accreted material? does it accumulate?
- is the material "incorporated" into the crust?
- is the accretion mound static - in "equilibrium"?
- can the accretion mound grow or shrink (height/total mass)?
- can the B-field configuration change? (e.g. "ballooning")
- does material "leak out" to the sides at the base of the accretion mound?
- how much total mass can be "stored" in the mound?
(until the B-field "breaks")
- can the B-field be "screened" or "buried"?
- what is the dynamical evolution of the mound?

Why does E_{cyc} increase with L_x ?

Why does E_{cyc} decrease with time?



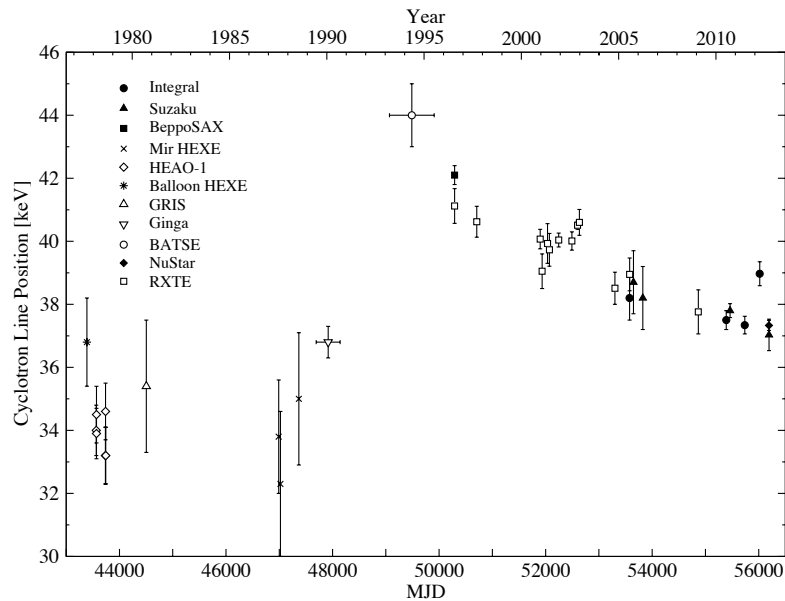
$$M_{\text{acc}} \sim 10^{-5} M_{\odot}$$

Payne & Melatos 2004
MNRAS 351, 569

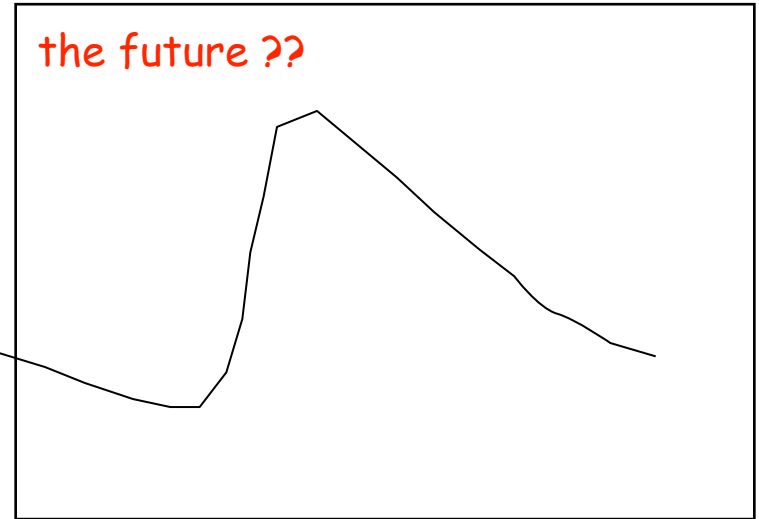
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Very-long-term behavior of E_{cyc} ??

Her X-1



the future ??



does E_{cyc} possibly follow a cyclic behavior on long time scales?

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